



Online Calibration of the WSRT Phased Array Feed system

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ASTRON is part of the Netherlands Organisation for Scientific Research (NWO) PAF Workshop, Cagliari, Italy, 24 August 2016

Outline



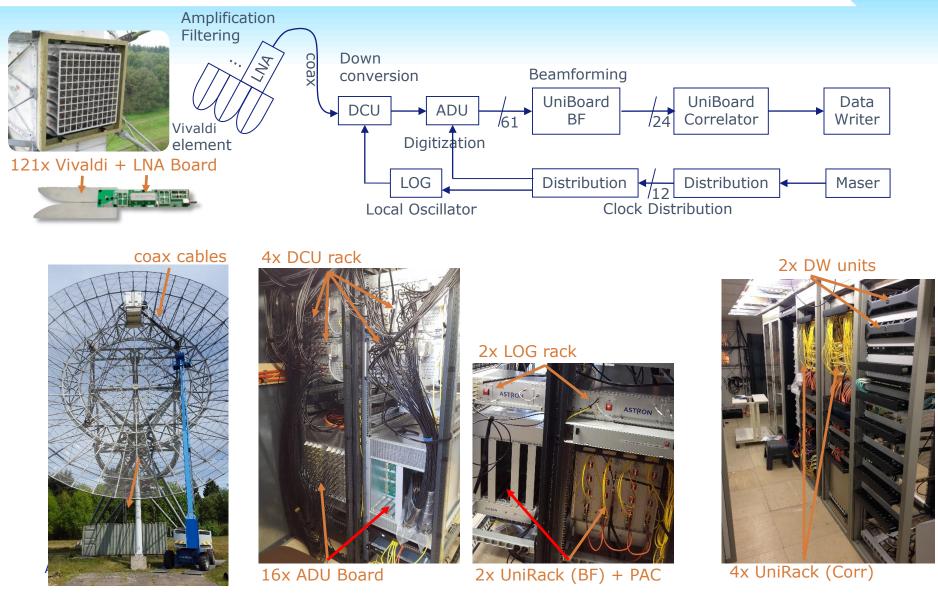


- Phased Array Technology
- Why calibration?
- How to calibrate?
 - Calibration with or without reference signal
- Online calibration scheme for APERTIF
 - Two designs
- Future plans

PAF Technology: many receiver chains













- Difference in gain and phase of individual elements
 - Sun radiates on some cables directly



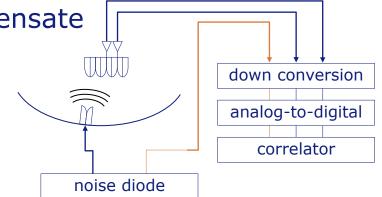
- Drift between receiver chains affects beamshape
 - Requirement on beamshape
 - Requirement on beamformer weights
 - 0.3 dB in amp; 2 degree in phase

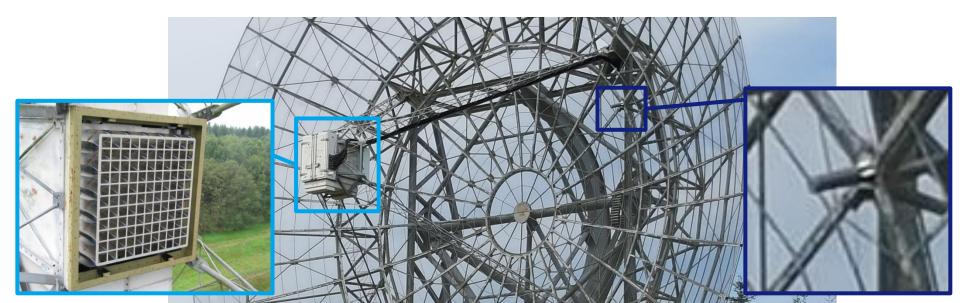
How to Calibrate? two options





- Emit stable broadband noise
- Track changes of this signal to compensate
- 1: Without reference signal (NxN)
- 2: With reference signal (Nx1)





option 1: without reference

• Record NxN correlation matrix R_{t}

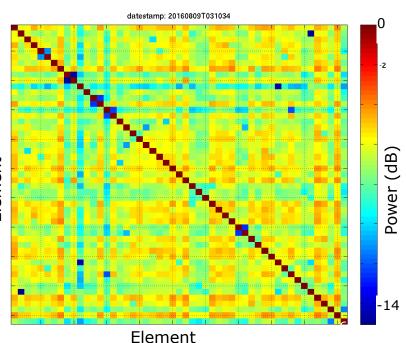
• At start $(t = t_0)$

How to Calibrate?

Every x seconds thereafter

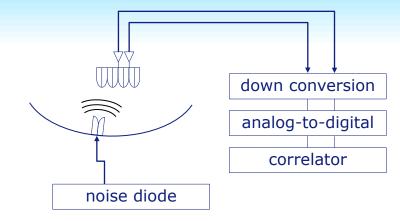
• Correction $\boldsymbol{w}_{corr_t} = \boldsymbol{e}_t \oslash \boldsymbol{e}_{t_0}$ e_t is the eigenvector, corresponding

co the largest eigenve.
⊘ denotes the element-wise division the element-wise division the element di divisio



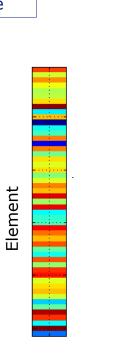


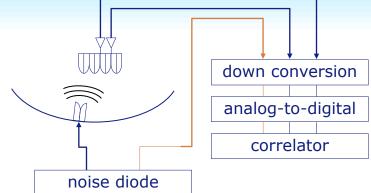
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How to Calibrate? option 2: with reference

- Correlating PAF elements against noise diode
- Record Nx1 correlation vector r_t
- Correction $w_{corr, t} = r_t \oslash r_{t_0}$
- Fewer correlation values can be traded for frequency bins
- Reduces time spending on onlinecalibration







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-2.5

Power (dB)

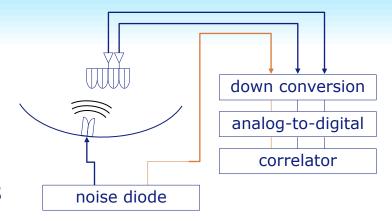
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Comparison Measurement



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- Measure full correlation matrix every second, for 1 hour
- Without reference, select correlation matrix of PAF elements
- With reference, select correlation vector with calsrc



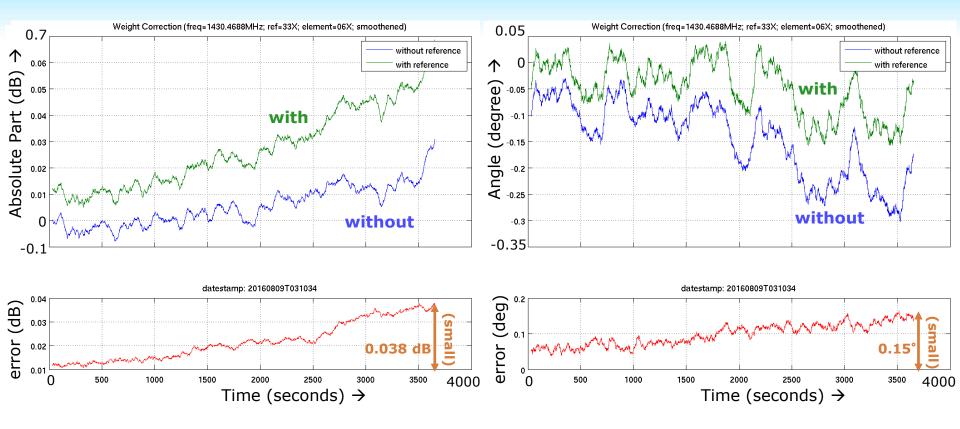
- Determine w_{corr t} for every second with and without reference
- Compare correction factor for a single PAF element
- Measurement during night time

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Result







- Adding reference gives similar result
 - Sky signal may cause variations

Result

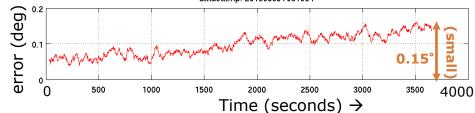


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Threshold for new calibration:

- 0.3 deg amp; 2 deg phase
- Threshold not reached within 1 hour!
- During night time, online calibration interval longer than 1 hour
- Adding reference gives same result





Sky signal may cause variations

Online Calibration: 6 Steps





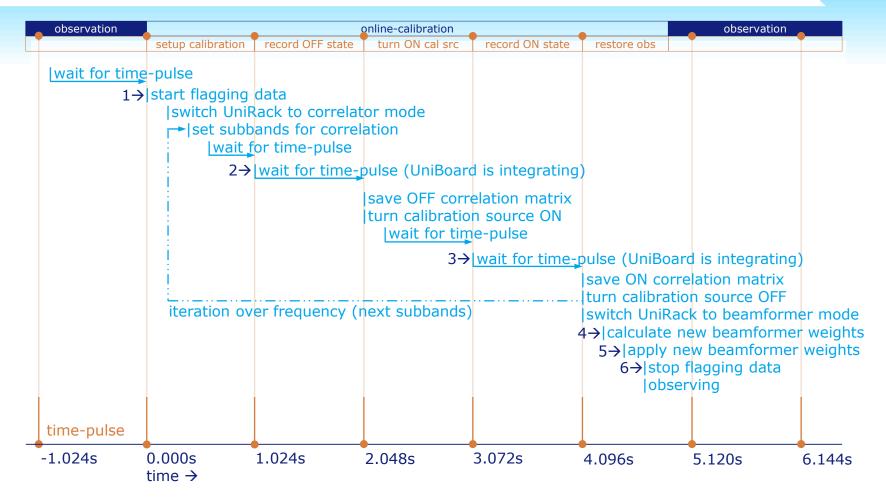
- 1. Flag data
- 2. Record **R**_{off}
- 3. Record R_{on} (or r_{on})
- 4. Calculate new weights: $\mathbf{R}_{on_t} \otimes \mathbf{R}_{on_t} \rightarrow \mathbf{b} \mathbf{f}$ weights
- 5. Apply new beamformer weights
- 6. Unflag data

R_{off} is included to track relative system temperature

Online Calibration: 6 Steps







total measurement time = (n_iterations * 4 + 1) * 1.024 sec

Online Calibration: two options



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A single loop records

- 64x64 complex correlation matrix for 6 subbands, or
- 1x64 complex correlation vector for 252 subbands

Option	Percentage of bandwidth per iteration	Percentage of bandwidth (total)	Number of iterations required	Time required for calibration measurement
64x64x6	1.5%	33%	22	91.136 s
		66%	44	181.248 s
(without	reference)	100%	67	275.456 s
1x64x252	66%	33%	1	5.12 s (17.8x faster)
		66%	1	5.12 s (35.4x faster)
(with reference)		100%	2	9.216 s (29.9x faster)

If it works, time improvement of factor >17.8

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Summary





- Two options for PAF calibration
- Adding a reference
 - Gives same correction factor
 - Reduces the time required for online calibration
 - by factor ~30
- Online calibration interval > 1 hour (system stable)
 - During night time
- Future
 - Test that varation between options is due to sky signal
 - Perform longer tests (including sunrise & sunset)